

**Erratum: “The First CHIME/FRB Fast Radio Burst Catalog” (2021, ApJS, 257, 59)**

The CHIME/FRB Collaboration, Mandana Amiri¹, Bridget C. Andersen^{2,3}, Kevin Bandura^{4,5}, Sabrina Berger^{2,3}, Mohit Bhargwa^{2,3}, Michelle M. Boyce⁶, P. J. Boyle^{2,3}, Charanjot Brar², Daniela Breitman^{7,8,9}, Tomas Cassanelli^{8,9}, Pragya Chawla^{2,3}, Tianyue Chen¹⁰, J.-F. Cliche², Amanda Cook^{8,9}, Davor Cubranic¹, Alice P. Curtin^{2,3}, Meiling Deng^{11,12}, Matt Dobbs^{2,3}, Fengqiu (Adam) Dong¹, Gwendolyn Eadie⁹, Mateus Fandino¹, Emmanuel Fonseca^{2,3}, B. M. Gaensler^{8,9}, Utkarsh Giri^{11,13}, Deborah C. Good¹, Mark Halpern¹, Alex S. Hill^{12,14}, Gary Hinshaw¹, Alexander Josephy^{2,3}, Jane F. Kaczmarek¹², Zarif Kader^{2,3}, Joseph W. Kania^{5,15}, Victoria M. Kaspi^{2,3}, T. L. Landecker¹², Dustin Lang^{11,13}, Calvin Leung^{10,16}, Dongzi Li^{7,17}, Hsiu-Hsien Lin¹⁷, Kiyoshi W. Masui^{10,16}, Ryan Mckinven^{8,9}, Juan Mena-Parra¹⁰, Marcus Merryfield^{2,3}, Bradley W. Meyers¹, Daniele Michilli^{2,3}, Nikola Milutinovic^{1,12}, Arash Mirhosseini¹, Moritz Münchmeyer^{11,18}, Arun Naidu^{2,3}, Laura Newburgh¹⁹, Cherry Ng⁸, Chitrang Patel^{2,8}, Ue-Li Pen^{7,8,9,11,17}, Emily Petroff^{2,3,20,25}, Tristan Pinsonneault-Marotte¹, Ziggy Pleunis^{2,3}, Masoud Rafiei-Ravandi^{11,13}, Mubdi Rahman^{8,21}, Scott M. Ransom²², Andre Renard⁸, Pranav Sanghavi^{4,5}, Paul Scholz⁸, J. Richard Shaw¹, Kaitlyn Shin^{10,16}, Seth R. Siegel^{2,3}, Andrew E. Sikora^{2,3}, Saurabh Singh², Kendrick M. Smith¹¹, Ingrid Stairs¹, Chia Min Tan^{2,3}, S. P. Tendulkar^{23,24}, Keith Vanderlinde^{8,9}, Haochen Wang^{10,16}, Dallas Wulf^{2,3}, and A. V. Zwaniga²

¹ Department of Physics and Astronomy, University of British Columbia, 6224 Agricultural Road, Vancouver, BC V6T 1Z1, Canada² Department of Physics, McGill University, 3600 rue University, Montréal, QC H3A 2T8, Canada³ McGill Space Institute, McGill University, 3550 rue University, Montréal, QC H3A 2A7, Canada⁴ Lane Department of Computer Science and Electrical Engineering, 1220 Evansdale Drive, PO Box 6109, Morgantown, WV 26506, USA⁵ Center for Gravitational Waves and Cosmology, West Virginia University, Chestnut Ridge Research Building, Morgantown, WV 26505, USA⁶ Department of Physics and Astronomy, University of Manitoba, Winnipeg, MB R3T 2N2, Canada⁷ Department of Physics, University of Toronto, 60 St. George Street, Toronto, ON M5S 1A7, Canada⁸ Dunlap Institute for Astronomy & Astrophysics, University of Toronto, 50 St. George Street, Toronto, ON M5S 3H4, Canada⁹ David A. Dunlap Department of Astronomy & Astrophysics, University of Toronto, 50 St. George Street, Toronto, ON M5S 3H4, Canada¹⁰ MIT Kavli Institute for Astrophysics and Space Research, Massachusetts Institute of Technology, 77 Massachusetts Ave., Cambridge, MA 02139, USA¹¹ Perimeter Institute for Theoretical Physics, 31 Caroline Street N, Waterloo, ON N2S 2YL, Canada¹² Dominion Radio Astrophysical Observatory, Herzberg Research Centre for Astronomy and Astrophysics, National Research Council Canada, PO Box 248, Penticton, BC V2A 6J9, Canada¹³ Department of Physics and Astronomy, University of Waterloo, Waterloo, ON N2L 3G1, Canada¹⁴ Department of Computer Science, Math, Physics, & Statistics, University of British Columbia, Kelowna, BC V1V 1V7, Canada¹⁵ Department of Physics and Astronomy, West Virginia University, PO Box 6315, Morgantown, WV 26506, USA¹⁶ Department of Physics, Massachusetts Institute of Technology, 77 Massachusetts Ave., Cambridge, MA 02139, USA; kshin@mit.edu¹⁷ Canadian Institute for Theoretical Astrophysics, 60 St. George Street, Toronto, ON M5S 3H8, Canada¹⁸ Department of Physics, University of Wisconsin–Madison, 1150 University Ave., Madison, WI 53706, USA¹⁹ Department of Physics, Yale University, New Haven, CT 06520, USA²⁰ Anton Pannekoek Institute for Astronomy, University of Amsterdam, Science Park 904, 1098 XH Amsterdam, The Netherlands²¹ Sidrat Research, PO Box 73527 RPO Wychwood, Toronto, ON M6C 4A7, Canada²² National Radio Astronomy Observatory, 520 Edgemont Rd., Charlottesville, VA 22903, USA²³ National Centre for Radio Astrophysics, Post Bag 3, Ganeshkhind, Pune, 411007, India²⁴ Department of Astronomy and Astrophysics, Tata Institute of Fundamental Research, Mumbai, 400005, India

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This erratum addresses two errors that occurred when inferring the CHIME/FRB all-sky rate of fast radio bursts (FRBs) above a fluence of 5 Jy ms at 600 MHz, with a scattering time at 600 MHz under 10 ms and dispersion measure (DM) above 100 pc cm^{−3}. Each error is just below the ≈20% level in the same direction, bringing the inferred all-sky rate down to $[525 \pm 30(\text{stat.})^{+142}_{-131}(\text{sys.})]$ bursts sky^{−1} day^{−1}.

In the published article, we obtain selection-bias-corrected FRB property distributions for fluence (F), DM, pulse width (w), and scattering timescale at 600 MHz (τ). Afterwards, in Section 6.2, we investigate the fluence distribution of FRBs, simultaneously inferring the all-sky rate.

The first error was a bug in the code used when normalizing the observation function. Instead of normalizing by a factor of the sum of the fiducial model weights, the observation function was normalized by a factor of the total number of injected FRBs. The second error stemmed from the fact that the fiducial model weights used in the analysis did not properly exclude bursts with a scattering timescale above 10 ms at 600 MHz, a population we found was poorly constrained with CHIME/FRB observations. Thus, while we quoted a rate of FRBs below a scattering timescale of 10 ms at 600 MHz, the previously quoted rate did extrapolate to that region.

²⁵ Veni Fellow.

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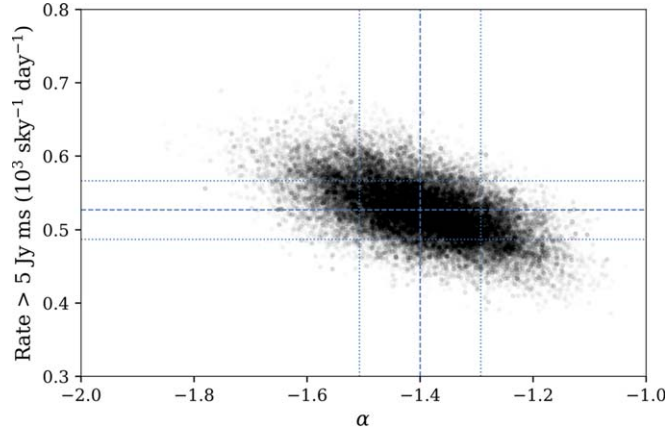


Figure 1. Markov Chain Monte Carlo samples of the posterior distributions for the all-sky rate and cumulative fluence distribution index α model parameters. The all-sky rate is for FRBs above 5 Jy ms, a DM above 100 pc cm^{-3} , and scattering timescales less than 10 ms at 600 MHz. Overlaid in dashed blue lines are the mean parameter values, and in dotted blue lines are the 1σ statistical uncertainty. Systematic uncertainties on the rate and α are not included in the distributions. The fluence distribution is still consistent with the Euclidean expectation of $\alpha = -3/2$.

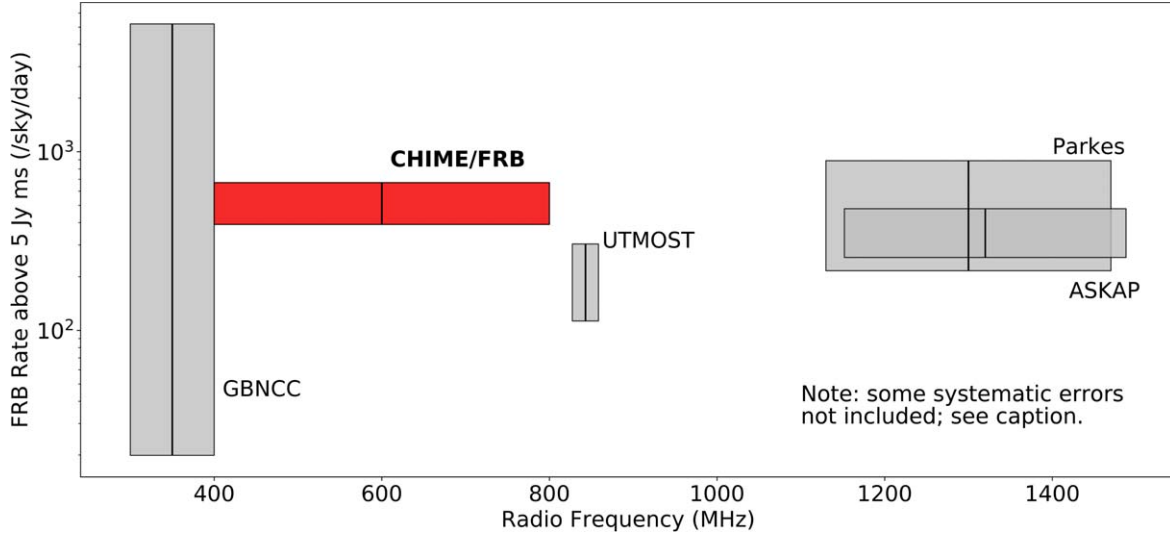


Figure 2. The updated CHIME/FRB all-sky rate plotted alongside other rate estimates. All rates are scaled (if needed) to a fluence limit of 5 Jy ms using our measured fluence index $\alpha = -1.40$, and plotted assuming a flat spectral index. Apparent discrepancies may be due to uncertainties and systematic differences from data being collected and processed by different surveys and instruments.

Above a fluence of 5 Jy ms at 600 MHz, with a scattering time at 600 MHz under 10 ms and DM above 100 pc cm^{-3} , the original quoted all-sky rate in the published article was $820 \pm 60(\text{stat.})^{+220}_{-200}(\text{sys.}) \text{ sky}^{-1} \text{ day}^{-1}$. After accounting for these two errors, neither of which affects our calculation of the systematic error budget, the updated all-sky rate of FRBs using the Catalog 1 sample is inferred to be $[525 \pm 30(\text{stat.})^{+142}_{-131}(\text{sys.})] \text{ sky}^{-1} \text{ day}^{-1}$. The index of the source counts distribution (i.e., the slope of $\log N - \log S$) remains unchanged to any statistical significance (Figure 1).

The updated sky rate measured from our Catalog 1 sample is $\approx 35\%$ lower than the erroneous sky rate measured. Thus, the interpretation of the rate is moderately, but not drastically, different than what was presented in Section 7.5 of the published article. The caveat of directly comparing our measured sky rate to others in the literature still holds due to the methods with which we accounted for instrumental effects. Nevertheless, we still present some simple comparisons. The CHIME/FRB rate is still consistent with the GBNCC rate of $3.4^{+15.4}_{-3.3} \times 10^3 \text{ sky}^{-1} \text{ day}^{-1}$ above ~ 2 Jy ms in the 300–400 MHz range (Parent et al. 2020). It is now also more consistent with the sky rate of $98^{+59}_{-39} \text{ sky}^{-1} \text{ day}^{-1}$ above a fluence of 8 Jy ms at 843 MHz reported from Molonglo/UTMOST (Farah et al. 2019), although one must still caution the interpretation of direct comparisons given the bandwidth difference between the two surveys. The updated rate is also still consistent with the sky rate of $1.7^{+1.5}_{-0.9} \times 10^3 \text{ sky}^{-1} \text{ day}^{-1}$ above a fluence limit of 2 Jy ms at 1.4 GHz from Parkes-detected FRBs (Bhandari et al. 2018), and more consistent with the ASKAP-derived rate of $37 \pm 8 \text{ sky}^{-1} \text{ day}^{-1}$ above 26 Jy ms at 1.4 GHz (Shannon et al. 2018) than it appeared before. These rates are shown in Figure 2 along with the updated CHIME/FRB all-sky rate.

The other results and conclusions in the published article are unaffected by these two errors. All other analyses contemporaneously presented are also unaffected by these errors (e.g., Josephy et al. 2021; Pleunis et al. 2021; Rafiei-Ravandi et al. 2021; Chawla et al. 2022). We look forward to further analyses and interpretations conducted using the first CHIME/FRB catalog, with the updated inferred all-sky rate.

ORCID iDs

Mandana Amiri  <https://orcid.org/0000-0001-6523-9029>
 Bridget C. Andersen  <https://orcid.org/0000-0001-5908-3152>
 Kevin Bandura  <https://orcid.org/0000-0003-3772-2798>
 Sabrina Berger  <https://orcid.org/0000-0002-4064-7883>
 Mohit Bhardwaj  <https://orcid.org/0000-0002-3615-3514>
 Michelle M. Boyce  <https://orcid.org/0000-0001-5470-3084>
 P. J. Boyle  <https://orcid.org/0000-0001-8537-9299>
 Charanjot Brar  <https://orcid.org/0000-0002-1800-8233>
 Daniela Breitman  <https://orcid.org/0000-0002-2349-3341>
 Tomas Cassanelli  <https://orcid.org/0000-0003-2047-5276>
 Pragya Chawla  <https://orcid.org/0000-0002-3426-7606>
 Tianyue Chen  <https://orcid.org/0000-0003-0173-6274>
 J.-F. Cliche  <https://orcid.org/0000-0001-6509-8430>
 Amanda Cook  <https://orcid.org/0000-0001-6422-8125>
 Davor Cubranic  <https://orcid.org/0000-0003-2319-9676>
 Alice P. Curtin  <https://orcid.org/0000-0002-8376-1563>
 Meiling Deng  <https://orcid.org/0000-0001-8123-7322>
 Matt Dobbs  <https://orcid.org/0000-0001-7166-6422>
 Fengqiu (Adam) Dong  <https://orcid.org/0000-0003-4098-5222>
 Gwendolyn Eadie  <https://orcid.org/0000-0003-3734-8177>
 Mateus Fandino  <https://orcid.org/0000-0002-6899-1176>
 Emmanuel Fonseca  <https://orcid.org/0000-0001-8384-5049>
 B. M. Gaensler  <https://orcid.org/0000-0002-3382-9558>
 Utkarsh Giri  <https://orcid.org/0000-0001-5553-9167>
 Deborah C. Good  <https://orcid.org/0000-0003-1884-348X>
 Mark Halpern  <https://orcid.org/0000-0002-1760-0868>
 Alex S. Hill  <https://orcid.org/0000-0001-7301-5666>
 Gary Hinshaw  <https://orcid.org/0000-0002-4241-8320>
 Alexander Josephy  <https://orcid.org/0000-0003-3059-6223>
 Jane F. Kaczmarek  <https://orcid.org/0000-0003-4810-7803>
 Zarif Kader  <https://orcid.org/0000-0003-2739-5869>
 Joseph W. Kania  <https://orcid.org/0000-0002-3354-3859>
 Victoria M. Kaspi  <https://orcid.org/0000-0001-9345-0307>
 T. L. Landecker  <https://orcid.org/0000-0003-1455-2546>
 Dustin Lang  <https://orcid.org/0000-0002-1172-0754>
 Calvin Leung  <https://orcid.org/0000-0002-4209-7408>
 Dongzi Li  <https://orcid.org/0000-0001-7931-0607>
 Hsiu-Hsien Lin  <https://orcid.org/0000-0001-7453-4273>
 Kiyoshi W. Masui  <https://orcid.org/0000-0002-4279-6946>
 Ryan Mckinven  <https://orcid.org/0000-0001-7348-6900>
 Juan Mena-Parra  <https://orcid.org/0000-0002-0772-9326>
 Marcus Merryfield  <https://orcid.org/0000-0003-2095-0380>

Bradley W. Meyers  <https://orcid.org/0000-0001-8845-1225>
 Daniele Michilli  <https://orcid.org/0000-0002-2551-7554>
 Nikola Milutinovic  <https://orcid.org/0000-0001-8292-0051>
 Arash Mirhosseini  <https://orcid.org/0000-0002-2626-5985>
 Moritz Münchmeyer  <https://orcid.org/0000-0002-3777-7791>
 Arun Naidu  <https://orcid.org/0000-0002-9225-9428>
 Laura Newburgh  <https://orcid.org/0000-0002-7333-5552>
 Cherry Ng  <https://orcid.org/0000-0002-3616-5160>
 Chitrang Patel  <https://orcid.org/0000-0003-3367-1073>
 Ue-Li Pen  <https://orcid.org/0000-0003-2155-9578>
 Emily Petroff  <https://orcid.org/0000-0002-9822-8008>
 Tristan Pinsonneault-Marotte  <https://orcid.org/0000-0002-9516-3245>
 Ziggy Pleunis  <https://orcid.org/0000-0002-4795-697X>
 Masoud Rafiei-Ravandi  <https://orcid.org/0000-0001-7694-6650>
 Mubdi Rahman  <https://orcid.org/0000-0003-1842-6096>
 Scott M. Ransom  <https://orcid.org/0000-0001-5799-9714>
 Andre Renard  <https://orcid.org/0000-0003-3463-7918>
 Pranav Sanghavi  <https://orcid.org/0000-0001-5504-229X>
 Paul Scholz  <https://orcid.org/0000-0002-7374-7119>
 J. Richard Shaw  <https://orcid.org/0000-0002-4543-4588>
 Kaitlyn Shin  <https://orcid.org/0000-0002-6823-2073>
 Seth R. Siegel  <https://orcid.org/0000-0003-2631-6217>
 Andrew E. Sikora  <https://orcid.org/0000-0002-1235-4485>
 Saurabh Singh  <https://orcid.org/0000-0001-7755-902X>
 Kendrick M. Smith  <https://orcid.org/0000-0002-2088-3125>
 Ingrid Stairs  <https://orcid.org/0000-0001-9784-8670>
 Chia Min Tan  <https://orcid.org/0000-0001-7509-0117>
 S. P. Tendulkar  <https://orcid.org/0000-0003-2548-2926>
 Keith Vanderlinde  <https://orcid.org/0000-0003-4535-9378>
 Haochen Wang  <https://orcid.org/0000-0002-1491-3738>
 Dallas Wulf  <https://orcid.org/0000-0001-7314-9496>
 A. V. Zwaniga  <https://orcid.org/0000-0001-8278-1936>

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