

Supernova Discoveries 2010–2011: Statistics and Trends

AVISHAY GAL-YAM

Department of Particle Physics and Astrophysics, Faculty of Physics, The Weizmann Institute of Science,
 Rehovot 76100, Israel; avishay.gal-yam@weizmann.ac.il; Co-chair,
 IAU Supernova Working Group, Kimmel Investigator

P. A. MAZZALI

Astrophysics Research Institute, Liverpool John Moores University, Liverpool, UK; INAF-Osservatorio Astronomico di Padova,
 vicolo dell'Osservatorio, 5, I-35122 Padova, Italy; Max-Planck Institut für Astrophysik, Karl-Schwarzschildstr. 1,
 D-85748 Garching, Germany; Co-chair, IAU Supernova Working Group

I. MANULIS

Department of Particle Physics and Astrophysics, Faculty of Physics, The Weizmann Institute of Science,
 Rehovot 76100, Israel

AND

D. BISHOP

Astronomy Section, Rochester Academy of Science, Rochester, NY; dbishop@vhd1.org

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ABSTRACT. We have inspected all supernova discoveries reported during 2010 and 2011, a total of 538 events during 2010 and 926 events during 2011. This number includes a small number of supernova impostors (bright extragalactic eruptions) but not novae or events that turned out to be Galactic stars. We examine the statistics of all discovered objects, as well as those of the subset of spectroscopically-confirmed events. In these 2 years we see the rise of wide-field non-targeted supernova surveys to prominence, with the largest numbers of events reported by the CRTS and PTF surveys (572 and 393 events in total respectively, contributing together 74% of all reported discoveries in 2011), followed by the integrated contribution of numerous amateurs (184 events). Among spectroscopically-confirmed events the PTF (393 events) leads, followed by CRTS (170 events), and amateur discoveries (144 events). Traditional galaxy-targeted surveys, such as LOSS and CHASE, maintain a strong contribution (86 and 61 events, respectively) with high spectroscopic completeness ($\sim 90\%$). It is interesting to note that the community managed to provide substantial spectroscopic follow-up for relatively brighter amateur discoveries ($\langle m \rangle = 16.5$ mag), but significant less help for fainter (and much more numerous) events promptly released by the CRTS ($\langle m \rangle = 18.6$ mag). Inspecting discovery magnitude and redshift distributions we find that PS1 discoveries have similar properties ($\langle m \rangle = 21.6$ mag, $\langle z \rangle = 0.23$) to events found in previous seasons by cosmology-oriented projects (e.g., SDSS-II), while PTF ($\langle m \rangle = 19.2$ mag, $\langle z \rangle = 0.095$) and CRTS ($\langle m \rangle = 18.6$ mag, $\langle z \rangle = 0.049$) populate the relatively unexplored phase space of faint supernovae (SNe) (> 19 mag) in nearby galaxies (mainly PTF), and events at $0.05 < z < 0.2$ (CRTS and PTF). Examining the specific question of reporting channels over the previous dozen years, we find that traditional reports via CBET telegrams now account only for a minority of SN discoveries.

1. INTRODUCTION

The study of supernovae (SNe) has a long history, going back to pioneers like Zwicky, Baade and Minkowski. Over the decades the numbers of SN discoveries reported every year have increased from a handful to hundreds. With improved instrumentation, fainter and more distant SNe are discovered. Here, we inspect the SN discoveries reported during the years 2010 and 2011 in detail. We also provide a report on the use of various reporting channels during the past decade. Our goal is to point

out general trends seen, some of which have operational implications to the community, in particular in allocation of follow-up resources and consideration of reporting mechanisms.

2. DATA

We have used the yearly summary provided by one of us (D. B.) via the “latest supernovae” webpage¹ to count the

¹ For more information see <http://www.rochesterastronomy.org/snimages/>.

TABLE 1
SUPERNOVA DISCOVERY STATISTICS

Project	Total number of SN discoveries (2010, 2011, total)	Spectroscopically confirmed events (2010, 2011, total)	Mean redshift (2010, 2011, total)	Mean magnitude (2010, 2011, total) (mag)
CRTS	190, 382, 572	67, 103, 170	0.057, 0.045, 0.049	18.0, 18.9, 18.6
PTF	88, 305, 393	88, 305, 393	0.086, 0.097, 0.095	19.3, 19.2, 19.2
Amateurs	82, 102, 184	69, 75, 144	0.019, 0.020, 0.020	16.7, 16.4, 16.5
LOSS	50, 36, 86	48, 28, 76	0.024, 0.020, 0.022	18.2, 17.3, 17.8
PS1	63, 11, 74	63, 11, 74	0.240, 0.169, 0.229	21.7, 20.8, 21.6
CHASE	36, 25, 61	32, 23, 55	0.023, 0.016, 0.020	16.6, 16.8, 16.7
MASTER	5, 23, 28	3, 9, 12	0.034, 0.159, 0.137	16.2, 17.6, 17.4
ROTSE	13, 6, 19	11, 5, 16	0.051, 0.061, 0.054	17.9, 17.6, 17.8
LSSSS	13, 6, 19	9, 3, 12	0.033, 0.035, 0.034	17.4, 16.8, 17.2
ISSP	0, 19, 19	0, 18, 18	0.022	16.8
LSQ	0, 7, 7	0, 7, 7	0.061	18.9
Others	11, 6, 17	9, 6, 15		

All explosive extragalactic objects reported are counted (including LBVs, LRNs and other SN impostors) with the exception of novae. All discoveries reported by individuals are counted as “amateur” discoveries; the “other” category includes all discoveries reported by projects with a few discoveries each. Project information: Catalina Real-Time Transient Survey (CRTS; <http://crtsc.caltech.edu/>; Drake et al. 2009; including subprojects MLS, SSS, SNhunt), Palomar Transient Factory (PTF; <http://www.astro.caltech.edu/ptf/>; Rau et al. 2009; Law et al. 2009), Pan-STARRs 1 (PS1; <http://ps1sc.org/transients/>), The Lick Observatory Supernova Search (LOSS; http://astro.berkeley.edu/bait/public_html/kait_lwd.html; Filippenko et al. 2001), The Chilean Automatic Supernova Search (CHASE; <http://www.das.uchile.cl/proyectoCHASE/>; Pignata et al. 2009), The ROTSE Supernova Verification Project (ROTSE-RSVP; <http://www.rotse.net/rsvp/>), The La Sagra Sky Survey (LSSS; <http://www.minorplanets.org/OLS/LSSS.html>), Mobile Astronomical System of Telescope Robots (MASTER; <http://observ.pereplet.ru/>; Lipunov et al. 2010), The Italian Supernova Search Project (ISSP; <http://italiansupernovae.org/>), The La-Silla Quest Variability Survey (LSQ; <http://hep.yale.edu/lasillaquest>). The mean magnitudes are based on values from the “Latest Supernovae” webpage, which are typically discovery magnitudes.

number of supernova discoveries during 2010–2011. As this page indexes discoveries reported via different channels (IAUCs, CBETs, ATEls, and the TOCP page) and attempts to cross-link objects reported by different projects at different times, it seems optimal for this purpose.² We have included in our statistics all objects initially reported as supernovae (including possible supernovae). We have separately counted spectroscopically-confirmed events. In principle, we did not count the few events that turned out to be supernova impostors, i.e., eruptions of Luminous Blue Variables (LBVs; e.g., Van Dyk 2005; Smith et al. 2011) or members of the class of Luminous Red Novae (LRN; Kulkarni et al. 2007; for example, PTF10fq, Kasliwal et al. 2011), and have eliminated reported events that turned out to be Galactic stars or extragalactic novae. A handful of such events (reported in 2010) may have been retained, but these should not influence our results. Objects reported independently by two projects were counted twice (so the total discovery number is slightly less than the sum of listed discovery numbers by project). The number of such “double-counting” events we identified is small (34 cases in 2011).

In general, the number of reported transients in 2011 was much larger than in 2010. In total there were 975 objects reported (551 in 2010). Of these, 49 (5.0%) were novae,

Luminous Blue Variables (LBVs), other extra-galactic variables, or Galactic variable stars. These are not covered by this paper.

We list in Table 1 the number of SN discoveries broken by projects. We have grouped all discoveries reported by individuals into an “amateur” class, and all discoveries by projects with a few SNe each into one “other” class. These data were manually harvested from the “latest Supernovae” webpage (accessed on 2010 December 31 and again during spring and summer 2012) and may contain errors either inherent to the webpage or introduced by us (if so, we apologize). Analysis of reports from 2011 was more complicated due to the larger number of events and additional reporting channels (e.g., private webpages) so we have tried to cross-check reported discoveries against the source of each report. We estimate that any remaining errors that may exist are negligible, and the trends seen in this summary should not be affected.

3. DISCUSSION

We scanned publically-accessible sources and have attempted to extract for as many events as we could the redshift of spectroscopically-confirmed SNe and their discovery magnitudes. The data were available for the majority of events. We calculated the mean values and present them in Table 1. In the few cases in which some data were missing (mostly redshifts), we calculated the means without these data; the incompleteness appears unlikely to change the reported numbers significantly. We did not calculate mean values for the “other”

² We have not included objects reported for the first time in refereed papers, e.g., those released by the PTF project in Arcavi et al. (2010) and Ofek et al. (2013a, 2013b), or those reported by PS1 in Chornock et al. (2013) and Berger et al. (2012).

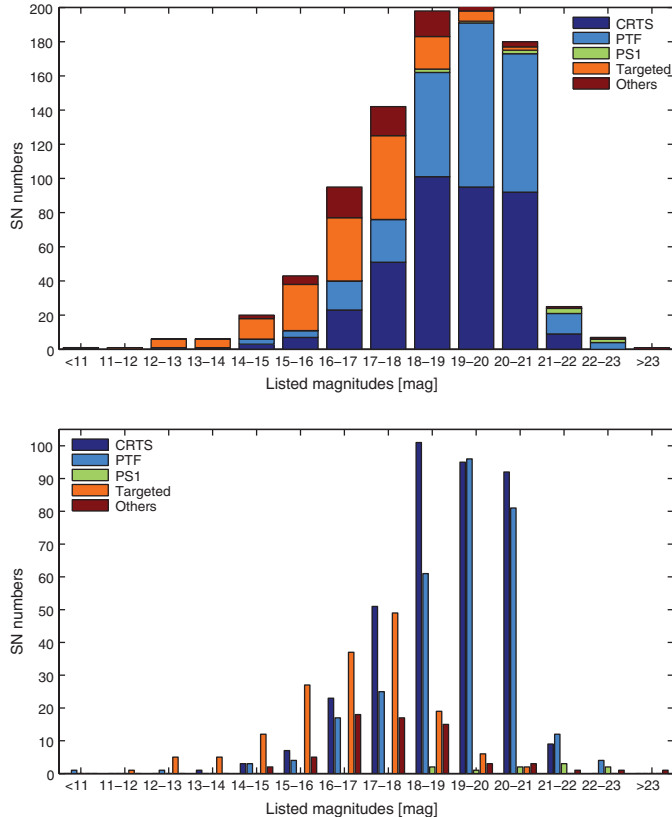


FIG. 1.—The magnitude distribution of SNe in 2011. Events between 18 and 21 mag dominate the numbers (*top*). Categorized by projects, the contributions of targeted surveys (Amateurs, LOSS and CHASE) remain dominant at brighter magnitudes, followed (moving to fainter objects), by CRTS, PTF and PS1.

category, as these SNe were discovered by very different means and projects.

Inspecting Table 1, we note the following trends.

The most prolific source of SN candidates in 2010–2011 was the CRTS survey. However, only $\sim 30\%$ of the candidates discovered by this survey have been spectroscopically confirmed, many by observers that appear to be affiliated with the survey. It is interesting to note that the larger SN community managed to provide substantial spectroscopic follow-up for relatively brighter amateur discoveries ($\langle m \rangle = 16.5$), but did not provide significant help for slightly fainter (and much more numerous) events promptly released by the CRTS ($\langle m \rangle = 18.6$). One can speculate that the amount of large telescope time set aside specifically for spectroscopic follow-up of SNe of unspecified origins (“public events”), in addition to flexible spectroscopic resources (nights during which observers can “squeeze in” recently discovered SNe) are saturated by the discovery stream from CRTS. In contrast, the LOSS survey that led this field in the previous decade (Leaman et al. 2011; Li et al. 2011) and which discovers SNe of similar or fainter magnitudes compared with those discovered by the CRTS, enjoyed high spectroscopic completeness ($\sim 90\%$), probably due to its dedicated

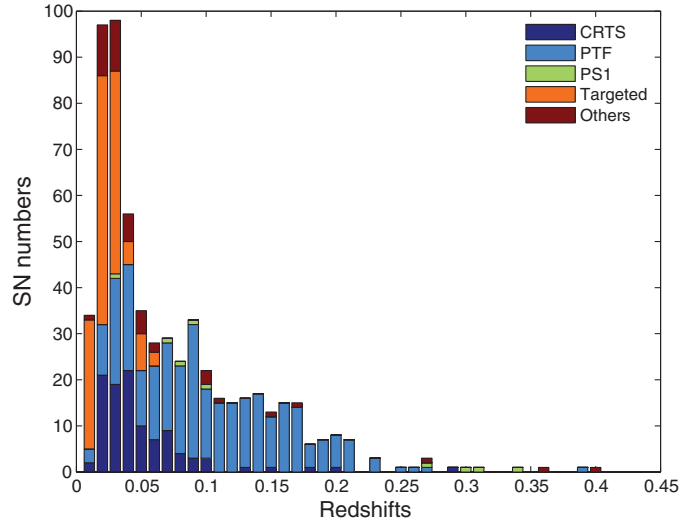


FIG. 2.—The redshift distribution of SNe discovered during 2011. Note the strong contribution by amateurs and targeted surveys at $z \sim 0.02$; followed by CRTS (up to $z \sim 0.1$), PTF (up to $z \sim 0.2$) and PS1 ($z > 0.25$).

long-standing follow-up program. The southern CHASE project typically discovered brighter events ($\langle m \rangle = 16.7$) and had a comparable spectroscopic completeness to that achieved for LOSS and amateur discoveries (which also have similar magnitudes). Since PTF and PS1 only report spectroscopically-confirmed events, their spectroscopic completeness cannot be determined.

Inspecting the discovery magnitude and redshift distributions (Figs. 1 and 2) we find that PS1 discoveries have similar properties ($\langle m \rangle = 21.6$, $\langle z \rangle \sim 0.23$) to events found in previous seasons by cosmology-oriented projects such as the SDSS-II supernova search (Frieman et al. 2008; Sako et al. 2008) and to a lesser degree also ESSENCE (Miknaitis et al. 2007) and SNLS (Astier et al. 2006).³ The two other new untargeted surveys, PTF ($\langle m \rangle = 19.4$, $\langle z \rangle = 0.099$) and CRTS ($\langle m \rangle = 18.6$, $\langle z \rangle \sim 0.049$), populate a relatively unexplored phase space. PTF is finding faint SNe (> 19 mag) in nearby galaxies, while both CRTS and PTF discover events in the redshift range $0.05 < z < 0.2$, previously little explored as it lies beyond the reach of the nearby targeted surveys (LOSS and CHASE) and yet requires a wide survey to discover numerous events (c.f., narrow and deep cosmological surveys).

In terms of reporting channels, in 2011 only 281 SNe (30.3%) were reported via CBET circulars, 580 SNe (62.6%)

³ The bulk of the PS1 discoveries which have been announced (in 2010) are from the PS1 Medium Deep Field survey, which covers a relatively narrow field (88 square degrees total) to deep limiting magnitudes ($m > 22$ mag), and hence provides fainter targets. A small number of SNe (and other transients) have been found (mostly in 2011) in the shallower, wider field “3Pi Faint Galaxy Supernova Survey” (e.g., Valenti et al. 2010), with discovery magnitudes in the range 17–20 mag.

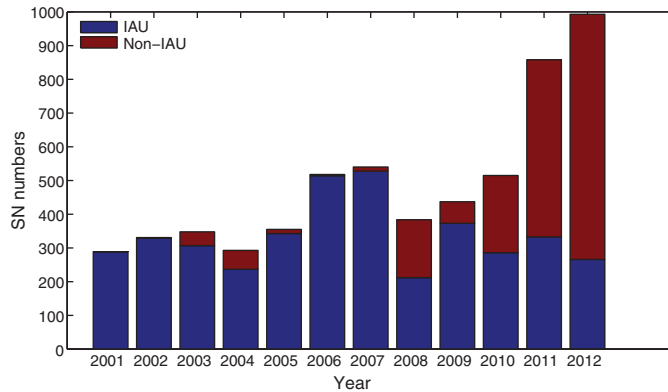


FIG. 3.—The numbers of SNe reported to the IAU (*blue*) or distributed via other channels (*red*). Starting in 2010, the numbers of non-IAU reports are significant, becoming the majority starting in 2011.

were reported in ATels, 57 SNe (6.2%) were reported only via the CBAT TOCP page, and 10 SNe (1.1%) were reported using the SNHunt project webpage. Some SNe were reported separately in both CBETs and ATels. We note that a change in CBET policy implemented in 2011 (reporting only spectroscopically-confirmed events) further complicates analysis and record-keeping (if one also includes SNe without spectroscopic confirmation). It may seem that a revision should be considered by the community regarding the mode in which SN discoveries are reported, tracked and circulated. Using statistics extracted

from the “latest supernovae” webpage for the years 2001–2012, we plot in Figure 3 the fractions of events reported to the IAU during this period. As can be seen, starting in 2010, a significant number of events are no longer reported to the IAU, and during 2011 (and more so, 2012) the non-IAU reports have become the majority.

4. SUMMARY

We enumerate the reported SN discoveries during 2010–2011 and point out several statistical trends. The impact of the new generation of untargeted surveys has been substantial during these years, with CRTS and PTF providing the largest number of the total and spectroscopically-confirmed events released to the community, respectively. The large number of CRTS SN candidates that have not been spectroscopically confirmed may indicate that the SN community cannot provide follow-up to the emerging large numbers of faint SNe. The inception and recent launch of the Public ESO Spectroscopic Survey of Transient Objects (PESSTO; PI Smartt⁴) may improve this situation. New surveys continue to provide access to new observational parameter space, including very faint events in nearby galaxies, and SNe in the relatively uncharted redshift range $0.05 < z < 0.2$.

⁴ For more information see <http://www.pessto.org>.

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